



Thermal conductivity of the disordered Fermi and electron liquids

Georg Schwiete

Johannes Gutenberg Universität Mainz

Alexander Finkel'stein

Texas A&M University, Weizmann Institute of Science,
and Landau Institute of Theoretical Physics

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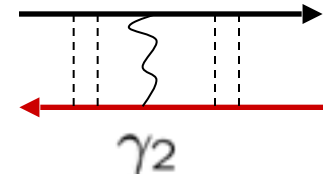
Perturbation theory – interacting systems

$$\delta\sigma(T) = -\frac{e^2}{\pi h} \left[1 + \left(1 - \frac{3}{2}\gamma_2 \right) \right] \log(1/T\tau)$$

WL correction
(Gang of 4)

1-singlet
contribution

3-triplet
contribution



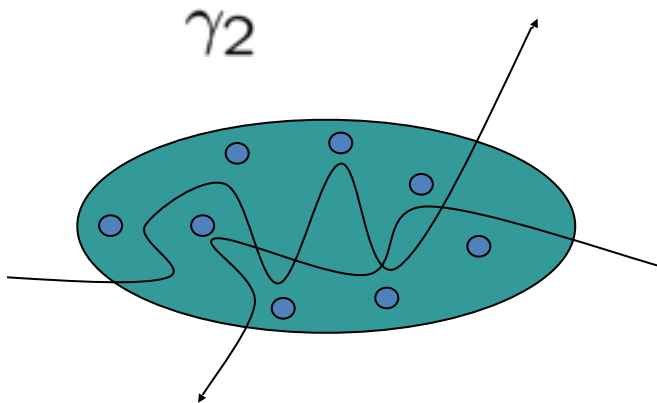
For $\gamma_2 \ll 1$, $\delta\sigma(T) < 0$ (insulating)

Altshuler, Aronov and Lee (1980), Finkel'stein (1983)

Disorder makes the
interaction scale-dependent

$$\delta\gamma_2(T) \sim \rho \log(1/T\tau) > 0$$

In the presence of disorder
 γ_2 is considerably enhanced at low T



Finkel'stein (1983)

Non-linear Sigma model (NL σ M)

Non-linear Sigma model: Effective low energy ($T < 1/\tau < E_F$) action for the disordered Fermi/electron liquid- Finkel'stein (1983)
 [noninteracting case: Wegner, Efetov Larkin Khmel'nitskii, ... (1979-)]

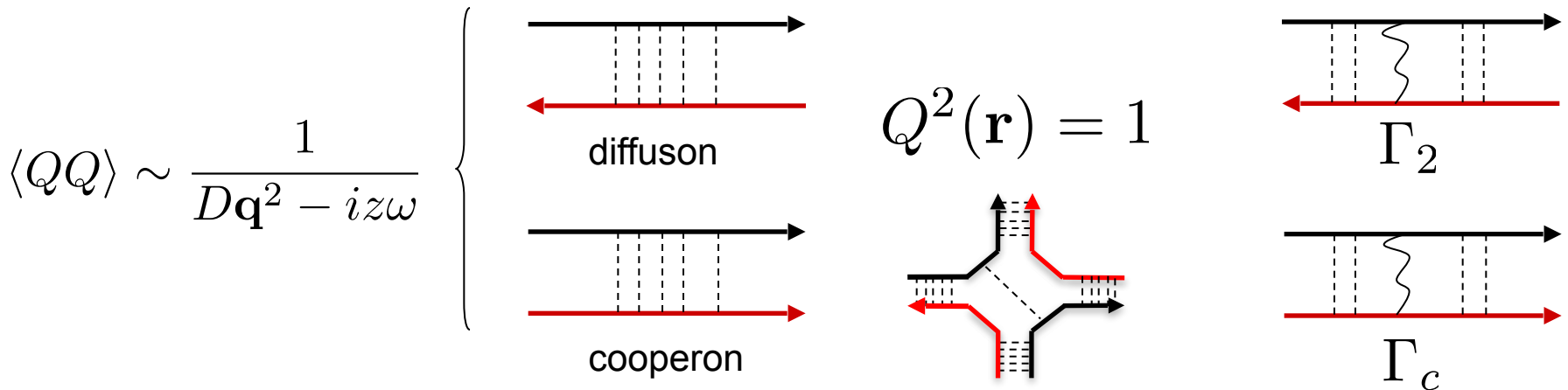
$$S[Q] \sim \int d\mathbf{r} \operatorname{tr} [D(\nabla Q)^2 + 4iz\hat{\epsilon}Q] + Q(\Gamma_1 + \Gamma_2 + \Gamma_c)Q$$

$$\rho \sim 1/D$$

frequency renormalization

$$\Gamma_2 = -\frac{F_0^\sigma}{1 + F_0^\sigma}$$

triplet-channel



Different methods: Replica/Keldysh

Structure of the RG equations

The interplay of disorder and interactions is captured by a set of coupled **Renormalization Group (scaling)** equations for ρ and γ_2

$$\frac{d \ln \rho}{d\xi} = \beta_\rho(\rho, \gamma_2)$$

$$\frac{d\gamma_2}{d\xi} = \beta_2(\rho, \gamma_2)$$

$$\xi = \ln(1/T\tau)$$

$$\gamma_2 = \frac{\Gamma_2}{z}$$

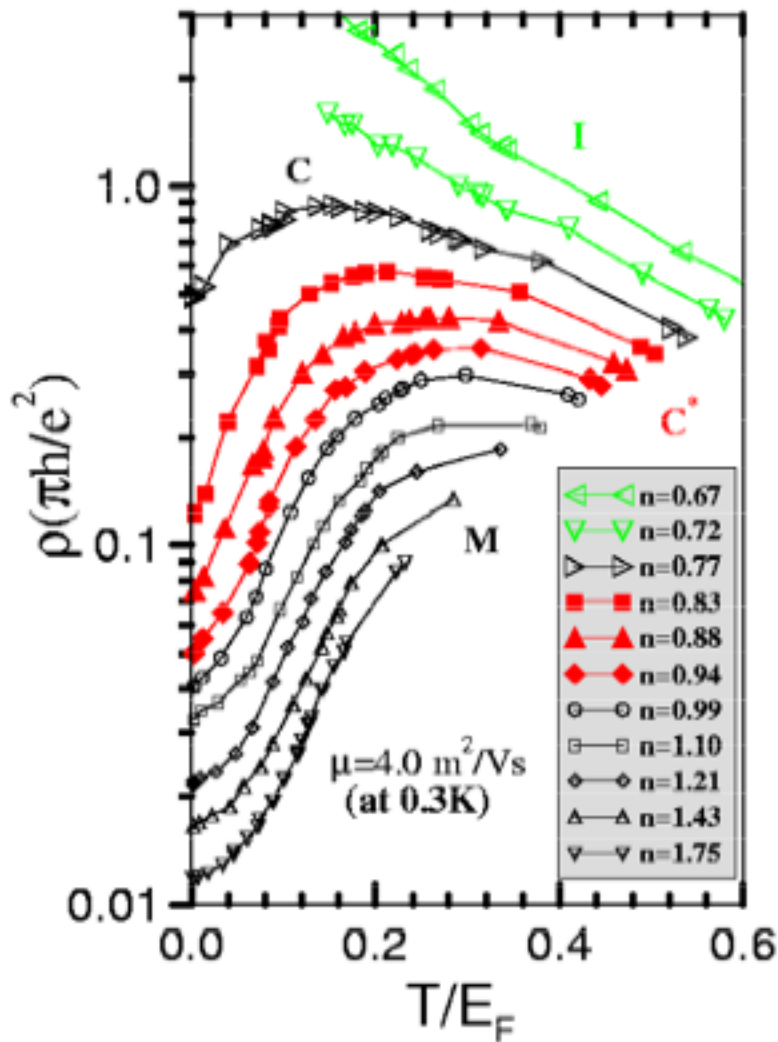
One more equation:

$$\frac{d \ln z}{d\xi} = \beta_z(\rho, \gamma_2)$$

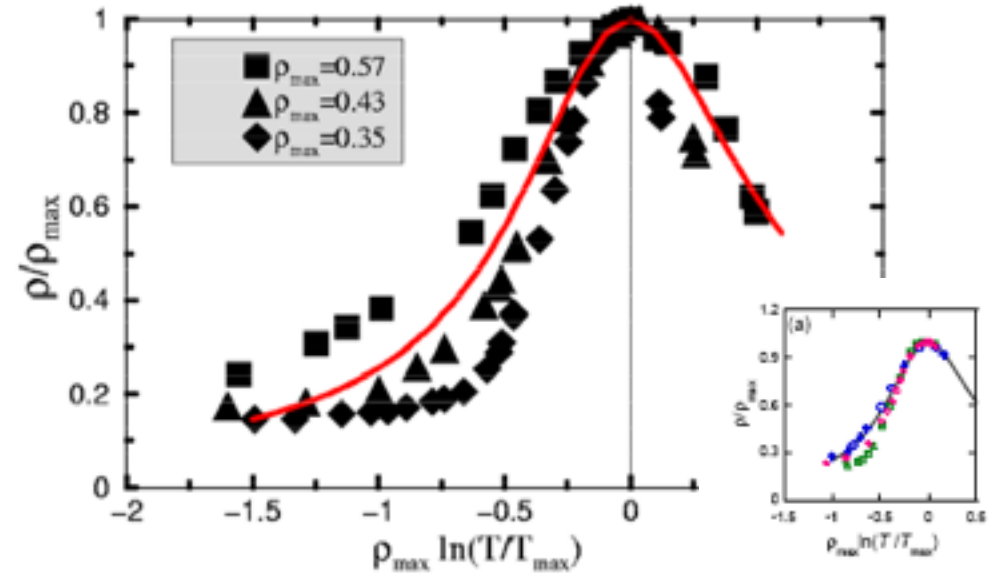
1-loop: leading order in ρ , all orders in the interaction.

does not affect the flow of ρ and γ_2 ,
important to understand thermodynamic properties

Analysis of high-mobility sample with RG for **two valleys**



Pudalov, et al., ('98)



Data from the region **C*** in a high-mobility sample. **No adjustable parameters** are used.

A. Punnoose and A. Finkelstein, PRL (2002)
S. Anissimova et al., Nature Physics (2007)

Thermal transport and the Wiedemann Franz law

Transport coefficients

$$\begin{pmatrix} \mathbf{j}_e \\ \mathbf{j}_k \end{pmatrix} = \begin{pmatrix} \sigma & \alpha\sigma \\ \Pi\sigma & \tilde{\kappa} \end{pmatrix} \begin{pmatrix} \mathbf{E} \\ -\nabla T \end{pmatrix}$$

Electric current

$$\mathbf{j}_e$$

Heat current

$$\mathbf{j}_k = \mathbf{j}_\varepsilon - \mu\mathbf{j}_n$$

Energy
current

Particle
current

Electric
conductivity

 σ
 α

Seebeck
coefficient

Onsager relation

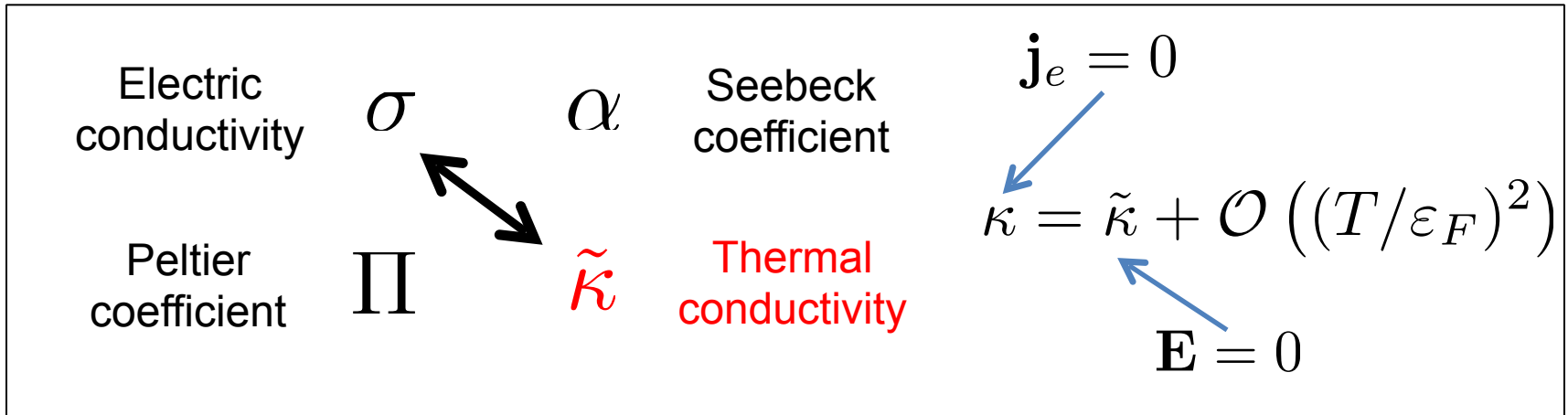
Peltier
coefficient

 Π
 $\tilde{\kappa}$

Thermal
conductivity

$$\Pi = T\alpha$$

The Wiedemann-Franz law



The Wiedemann-Franz “law”

$$\kappa = \mathcal{L}_0 \sigma T$$

Lorenz number

$$\mathcal{L}_0 = \frac{\pi^2}{3} \frac{k_B^2}{e^2} = \frac{c_{FL}}{2\nu e^2 T}$$

The Wiedemann Franz law is an approximate low-temperature relation for itinerant electron systems. What is the range of validity?

Heat transport and the Wiedemann-Franz law in disordered electron systems - History of the problem

- **Wiedemann-Franz law ($\kappa/\sigma T = \text{const.}$) holds**
for noninteracting disordered electron systems - *Chester, Thellung (1961)*.

- **Wiedemann-Franz law holds**
for a Fermi liquid - *Langer (1962)*.

After the development of the scaling theory of localization for interacting electrons
[Finkel'stein 83, Castellani et al. 84]:

- **Wiedemann-Franz law holds**
for the disordered electron liquid (renormalized perturbation theory, Ward Identities)
- *Castellani, di Castro, Kotliar, Lee, Strinati (1987-)*.

- **Wiedemann-Franz law violated**
for the disordered electron liquid (perturbation theory)

Kubo-formula – *Arfi (1992), Niven, Smith (2005)*.

Kinetic equation approaches - *Livanov et al. (1991), Raimondi et al. (2004), Catelani, Aleiner (2005), Michaeli, Finkelstein (2009)*.

While approaches differ, the result is common: Additional corrections, Wiedemann-Franz law violated

Can one resolve the contradiction and construct a comprehensive theory (including RG and additional log corrections) ?

Can one generalize the RG approach to thermal transport?

- **How is heat transported through the system?**
- **What are the consequences of replacing the electric field by a temperature gradient**

How to approach the problem?

How to do RG including a temperature gradient?

Perturbative calculations for κ were (mostly) based on [kinetic equation approaches](#).
Including a temperature gradient is straightforward, but how to do RG?

The scaling theory for σ was developed on the basis of a field theory (Nl σ M) with source fields. How to account for a temperature gradient?

$$\text{Nl}\sigma\text{M} \xrightarrow{\text{source } \varphi} \langle nn \rangle \xrightarrow{\text{Einstein}} \sigma$$

Our approach: Renormalize the Nl σ M with source fields
(Luttinger's „**gravitational potential**“ mimics temperature variation).


$$\text{Nl}\sigma\text{M} \xrightarrow{\text{source } ??} \langle kk \rangle \xrightarrow{\text{Einstein}} \kappa$$

Source fields for the heat density correlation function

Action:

$$S[\psi^*, \psi] = \int_{\mathbf{r}, t} (\psi^* i \partial_t \psi - k[\psi^*, \psi])$$

$$\mathcal{Z} = \int D(\psi, \psi^*) e^{iS} \quad k = h_0 + h_{int} - \mu n$$



$$S[\psi^*, \psi] = \int_{\mathbf{r}, t} (\psi^* i \partial_t \psi - (1 + \eta) k[\psi^*, \psi])$$

$$\chi_{kk} = \frac{i}{2} \frac{\delta^2 \mathcal{Z}}{\delta \eta_{\mathbf{r}_1 t_1} \delta \eta_{\mathbf{r}_2 t_2}}$$

Gravitational potential

Source fields for the heat density correlation function

$$S[\psi^*, \psi] = \int_{\mathbf{r}, t} (\psi^* i \partial_t \psi - (1 + \eta) k[\psi^*, \psi])$$

Problem: $S_{dis} = - \int_{\mathbf{r}, t} (1 + \eta) \psi^* u_{dis} \psi$

Change of variables: $\psi \rightarrow \frac{1}{\sqrt{1 + \eta}} \psi$ $\psi^* \rightarrow \psi^* \frac{1}{\sqrt{1 + \eta}}$

After this transformation, the derivation of the NL σ M is straightforward:

$$S[Q] \sim \int d\mathbf{r} \text{tr} [D(\nabla Q)^2 + 2iz\{\hat{\epsilon}, \lambda\}Q] + Q\lambda(\Gamma_1 + \Gamma_2)Q$$

$$\lambda = \frac{1}{1 + \eta} \approx 1 - \eta + \eta^2 + \dots$$

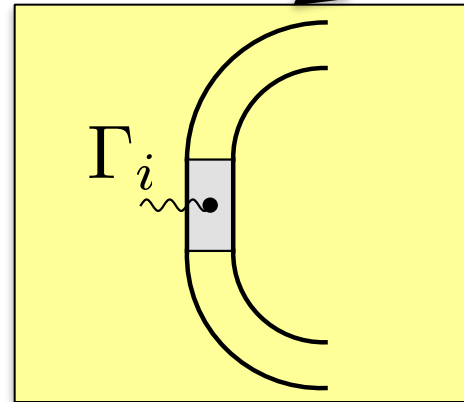
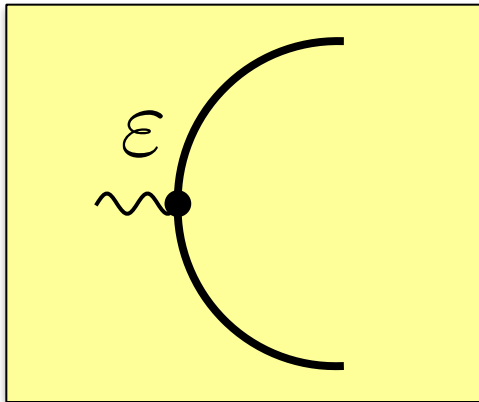
nonlinear in η !

NI σ M with “gravitational potentials”

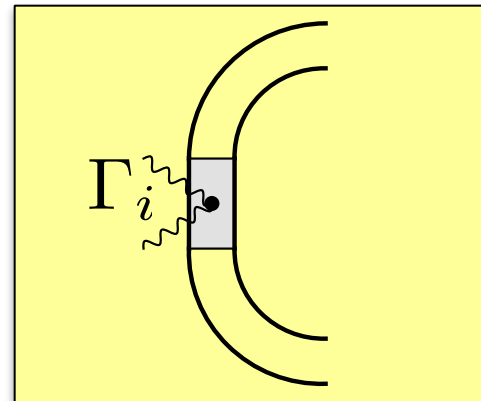
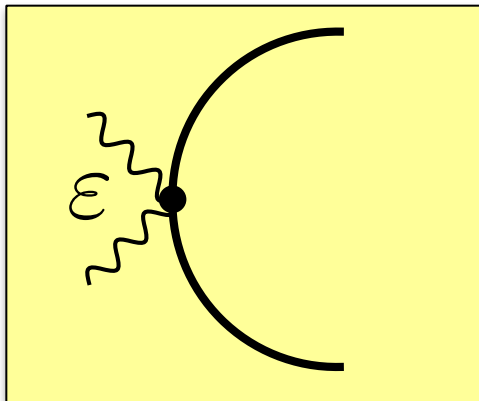
$$S[Q] \sim \int d\mathbf{r} \operatorname{tr} [D(\nabla Q)^2 + 2iz\{\hat{\epsilon}, \lambda\}Q] + Q\lambda(\Gamma_1 + \Gamma_2)Q$$

$$\lambda \approx 1 - \eta + \eta^2$$

η



η^2



The correlation function

Heat density correlation function
in the diffusive limit

$$\chi_{kk} = -cT \frac{D_k \mathbf{q}^2}{D_k \mathbf{q}^2 - i\omega}$$

Static limit

$$\chi_{kk}(\mathbf{q} \rightarrow 0, \omega = 0) = -cT$$

Conservation law

$$\chi_{kk}(\mathbf{q} = 0, \omega \rightarrow 0) = 0$$

Thermal conductivity

$$\kappa = cD_k$$

Specific heat

1. Direct calculation: Heat density - specific heat
(linear terms in η)
2. Static part of the correlation function
(quadratic terms in η)

Heat density and specific heat

$$S[Q] \sim \int dr \text{tr} [D(\nabla Q)^2 - 2z\{\hat{\varepsilon}, 1 - \eta + \eta^2\}Q] + Q(1 - \eta + \eta^2)(\Gamma_1 + \Gamma_2)Q$$

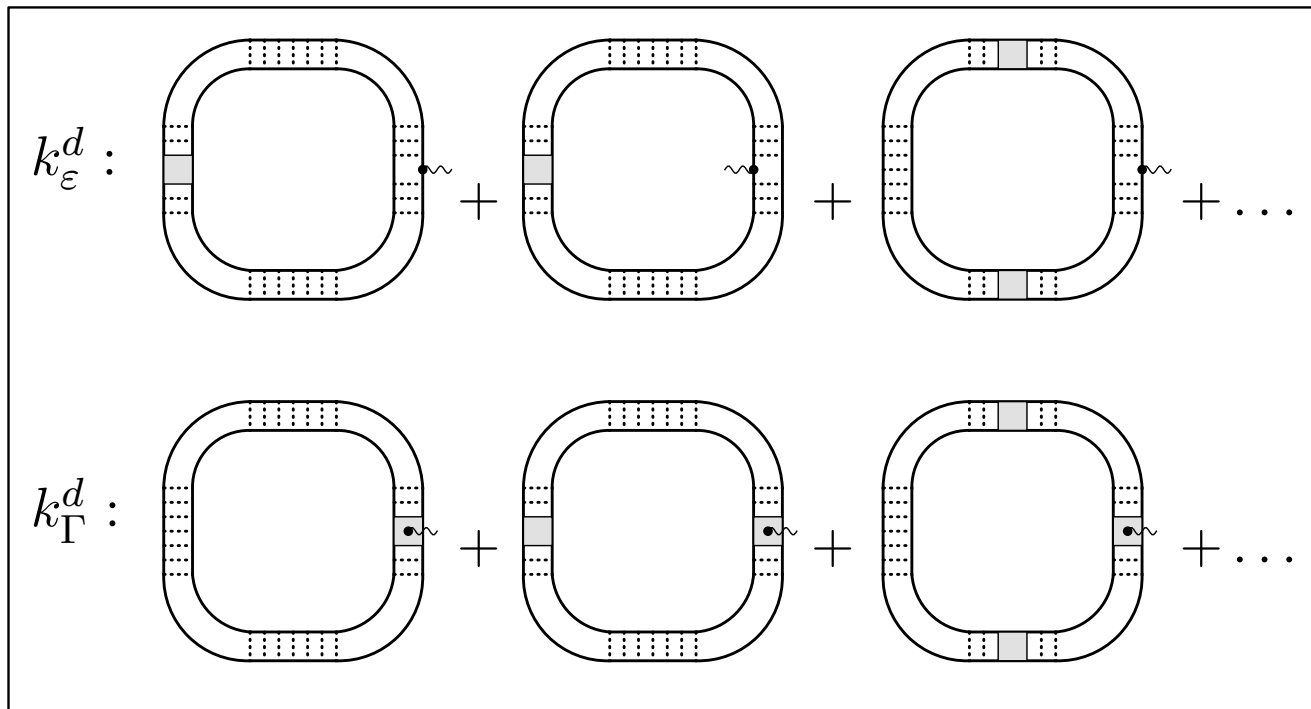
Heat density

$$k_{\eta=0}^d = \frac{i}{2} \frac{\delta \mathcal{Z}}{\delta \eta} \Big|_{\eta=0}$$

Specific heat

$$\delta c = \partial_T k_{\eta=0}^d = \delta z c_{FL}$$

Castellani, Di Castro (1986)

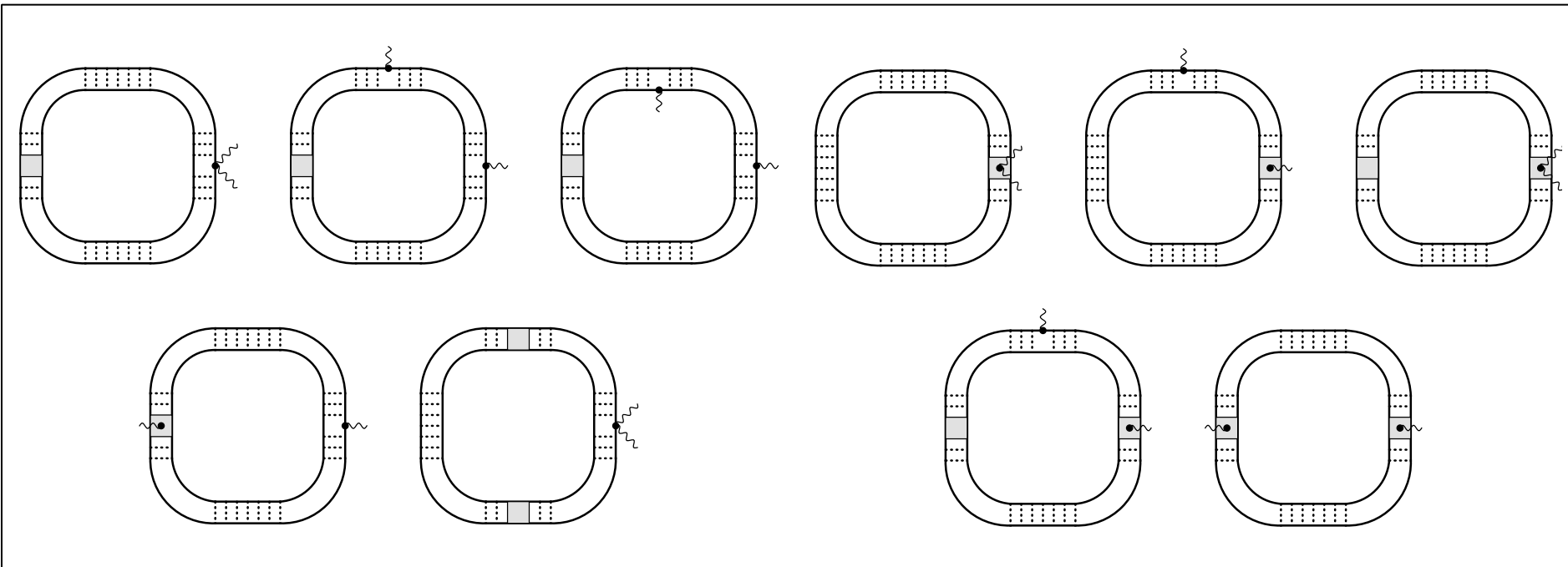


Specific heat and the static part of the correlation function

$$S[Q] \sim \int dr \text{tr} [D(\nabla Q)^2 - 2z\{\varepsilon, 1 - \eta + \eta^2\}Q] + Q(1 - \eta + \eta^2)(\Gamma_1 + \Gamma_2)Q$$

$$\chi_{kk}^{st} = \frac{i}{2} \frac{\delta^2 \mathcal{Z}}{\delta \eta^2} = -Tc$$

$$c = zC_{FL}$$



RG and the dynamical part of the correlation function

RG and the dynamic part of the correlation function

$$S = \int \text{tr}[D(1 + \zeta_D)(\nabla Q)^2 + 2iz\{\hat{\varepsilon}, 1 + \zeta_z\}Q] + \sum_{i=1,2} Q(1 + \zeta_{\Gamma_i})\Gamma_i Q$$

Initial conditions: $\zeta_D = 0 \quad \zeta_z = \zeta_{\Gamma_1} = \zeta_{\Gamma_2} = -\eta$

Parameterization: $Q = u U_s Q_f U_s^{-1} u^{-1} \quad Q^2 = 1$

$$Q_f = U_f \Lambda U_f^{-1}$$

fast

$$Q_s = U_s \Lambda U_s^{-1}$$

slow

u
slowest: distribution function

$$U_{\varepsilon_1 \varepsilon_2}^{-1} \zeta_i(\varepsilon_2 - \varepsilon_3) U_{\varepsilon_3 \varepsilon_4} \neq \zeta_i(\varepsilon_1 - \varepsilon_4)$$

$$\Delta(D\zeta_D) = \left(\zeta_D D \frac{\partial}{\partial D} + \zeta_z z \frac{\partial}{\partial z} + \zeta_{\Gamma_1} \Gamma_1 \frac{\partial}{\partial \Gamma_1} + \zeta_{\Gamma_2} \Gamma_2 \frac{\partial}{\partial \Gamma_2} \right) \Delta D$$

$$\Delta(z\zeta_z) = \left(\zeta_D D \frac{\partial}{\partial D} + \zeta_z z \frac{\partial}{\partial z} + \zeta_{\Gamma_1} \Gamma_1 \frac{\partial}{\partial \Gamma_1} + \zeta_{\Gamma_2} \Gamma_2 \frac{\partial}{\partial \Gamma_2} \right) \Delta z$$

⋮

RG and the dynamic part of the correlation function

$$S = \int \text{tr}[D(1 + \zeta_D)(\nabla Q)^2 + 2iz\{\hat{\epsilon}, 1 + \zeta_z\}Q] + \sum_{i=1,2} Q(1 + \zeta_{\Gamma_i})\Gamma_i Q$$

Result: Fixed point

$$\Delta\zeta_D = \Delta\zeta_z = \Delta\zeta_{\Gamma_1} = \Delta\zeta_{\Gamma_2} = 0$$

$$\zeta_D = 0 \quad \zeta_z = \zeta_{\Gamma_1} = \zeta_{\Gamma_2} = -\eta$$

$$S = \int \text{tr}[D(\nabla Q)^2 + 2iz\{\hat{\epsilon}, 1 - \eta\}Q] + Q(1 - \eta)(\Gamma_1 + \Gamma_2)Q$$

$$\chi_{kk} = -cT \frac{D_k \mathbf{q}^2}{D_k \mathbf{q}^2 - i\omega} \quad D_k = \frac{D}{z}$$

Conductivities and the Wiedemann Franz law

(Generalized) Einstein relations:

$$\sigma = e^2 \lim_{\omega \rightarrow 0} \lim_{q \rightarrow 0} \frac{\omega}{q^2} \text{Im} \chi_{nn}(\mathbf{q}, \omega) = e^2 \frac{\partial n}{\partial \mu} D_n = 2\nu e^2 D$$

$$\kappa = -\frac{1}{T} \lim_{\omega \rightarrow 0} \lim_{q \rightarrow 0} \frac{\omega}{q^2} \text{Im} \chi_{kk}(\mathbf{q}, \omega) = c D_k = c_{FL} D$$

$$D_n = \frac{D}{\frac{\partial n}{\partial \mu} / 2\nu} \quad D_k = \frac{D}{z} = \frac{D}{c/c_{FL}}$$

$$\kappa = \frac{c_{FL}}{2\nu e^2} \sigma$$

The structure immediately implies:

**Wiedemann Franz law is not violated within the RG regime ($T < \varepsilon < 1/\tau$),
neither for short-range nor for long-range (Coulomb) interaction.**

Something is missing in this treatment!

Beyond RG – the low temperature regime

General structure of the correlation functions

Density-density correlation function:

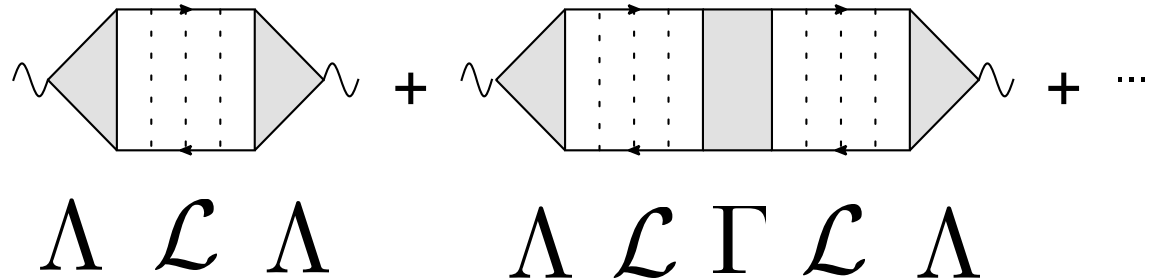
$$\chi_{nn} = \frac{dn}{d\mu} + 2\nu\Lambda_n^2 \frac{i\omega}{\mathcal{L}^{-1} + 2i\tilde{\Gamma}_s\omega}$$

Heat density-heat density correlation function:

$$\chi_{kk} = -cT - c_{FL}T\Lambda_k^2 \frac{i\omega}{\mathcal{L}^{-1}}$$

Castellani et al. (1987)

Dynamical part:

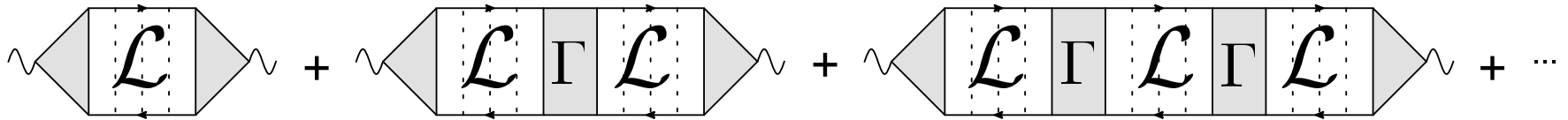


$$\mathcal{L} = \frac{\xi^2}{Dq^2 - iz\omega}$$

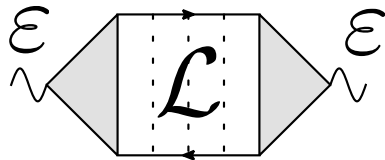
Renormalization of $D, \xi, z, \Lambda, \Gamma$

Why heat transport is different

χ_{nn}^{dyn} :

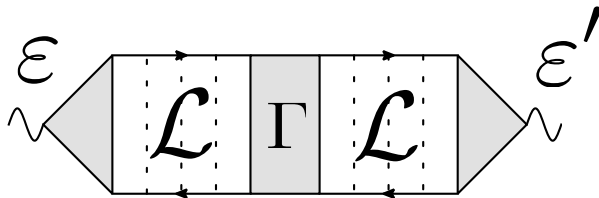


χ_{kk}^{dyn} :



$$\varepsilon^2 \rightarrow T^2$$

WF



$$\varepsilon \Gamma \varepsilon' \rightarrow 0$$

Rescattering is suppressed for constant Γ !

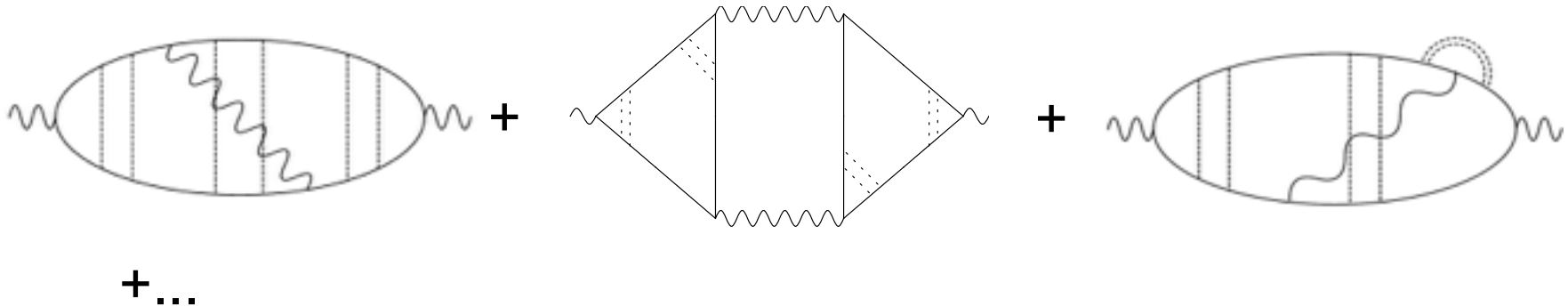
Finite result only for $\Gamma = \Gamma(\varepsilon, \varepsilon')$

Such contributions are not part of the traditional RG scheme!

Additional logarithms

For **short-range interactions** there are no additional (logarithmic) corrections.

For the **electron liquid (Coulomb interaction)** there are **additional logarithmic contributions** from scattering processes with sub-T frequency transfer.



All contributions are proportional to $\text{Im}(U^R)$: Decay into particle-hole pairs.

Example:

$$\delta\chi_{kk} \propto \int_{\mathbf{k}, \varepsilon, \nu} \varepsilon \nu \partial_{\varepsilon} F_{\varepsilon} (F_{\varepsilon+\nu} + F_{\varepsilon-\nu}) \text{Re} \mathcal{D}^2(\mathbf{k}, \nu) \text{Im} U^R(\mathbf{k}, \nu)$$

Corrections to heat conductivity

Additional logarithmic correction (not related to c):

$$\chi_{kk} = -cT \frac{D_k \mathbf{q}^2}{D_k \mathbf{q}^2 - i\omega}$$
$$D_k = \frac{1}{z} (D_n + \delta D^h)$$

Consistent with conservation law!

Additional logarithmic correction to κ :

$$\delta\kappa = \frac{T}{12} \log \frac{D\kappa_s^2}{T}$$

κ_s : screening radius

WF law is violated!

Agrees with the result of (recent) kinetic equation approaches

From the regime:

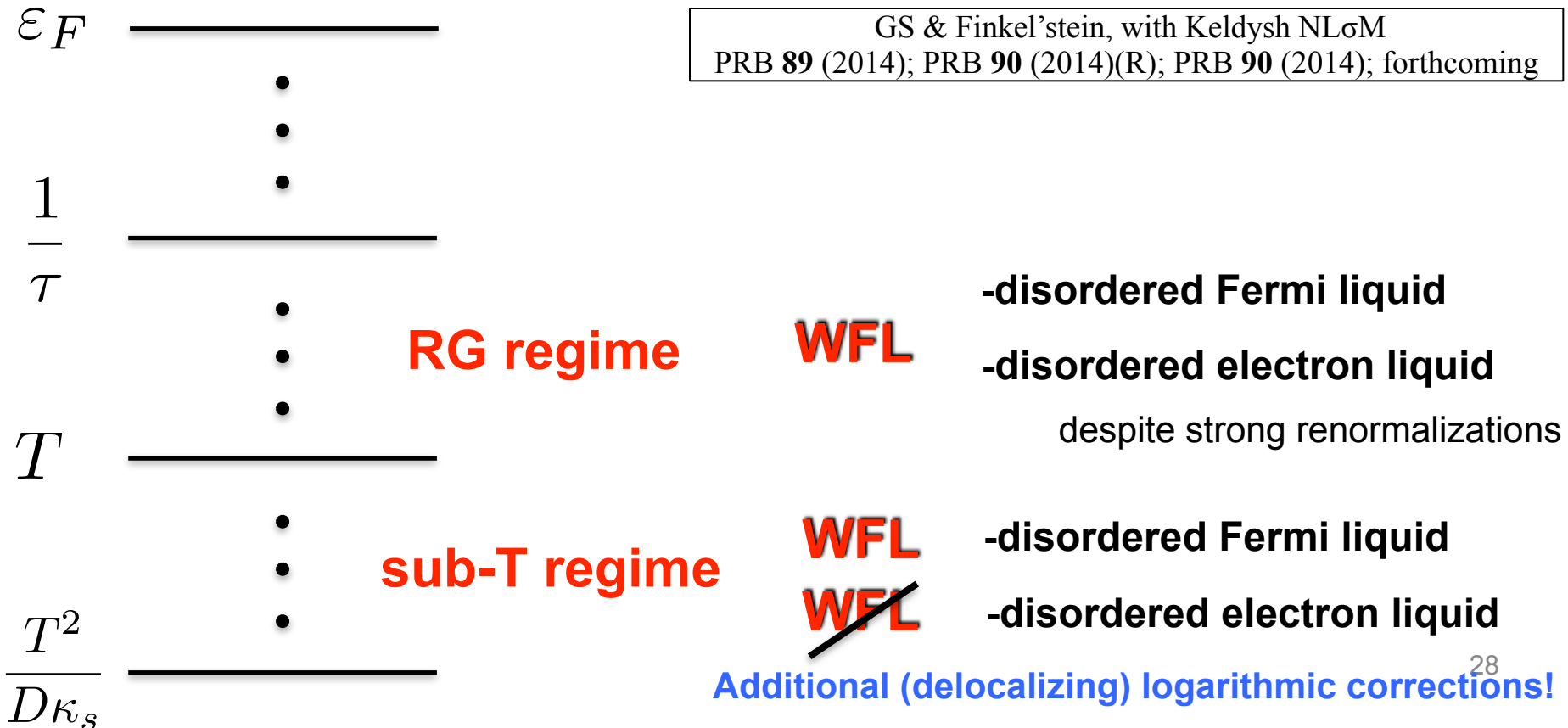
$$\frac{T^2}{D\kappa^2} < D\mathbf{k}^2 < T$$

Results: Thermal transport and the WFL

$$S[Q] \sim \int d\mathbf{r} \operatorname{tr}[D(\nabla Q)^2 + 2iz\{\hat{\varepsilon}, \lambda\}Q] + Q\lambda(\Gamma_1 + \Gamma_2)Q$$

Energy scales

$$\lambda \approx 1 - \eta + \eta^2$$



GS & Finkel'stein, with Keldysh NLσM
 PRB 89 (2014); PRB 90 (2014)(R); PRB 90 (2014); forthcoming

Summary

- We developed a field theoretic model with „gravitational potentials“ suitable for the analysis of heat density correlation function in the disordered electron liquid.
- The terms linear and quadratic in the gravitational potentials are consistent with each other and also with thermodynamics.
- For short range interactions the renormalization of κ and of σ are linked through the WF law.
- For long-range (Coulomb) interaction there are additional logarithmic corrections originating from outside of the RG regime. They lead to a violation of the WF law.

Georg Schwiete & Alexander Finkel'stein,
Phys. Rev. B **89** (2014); RG with Keldysh NL σ M;
Phys. Rev. B **90** (2014)(R); Wiedemann Franz law
Phys. Rev. B **90** (2014); RG for Keldysh NL σ M with grav. potentials
forthcoming: Analysis of low temperature regime

Thank you!